THEF-NIJM 89.02

# Phase shift analysis of all proton-proton scattering data below $T_{\rm lab}$=350 MeV*

J.R. Bergervoet, P.C. van Campen, R.A.M. Klomp, J.-L. de Kok, T.A. Rijken, V.G.J. Stoks, and J.J. de Swart

*Institute for Theoretical Physics, University of Nijmegen, Nijmegen, The Netherlands*

## Abstract

As a continuation of our 0–30 MeV analysis we present a multienergy phase shift analysis of all *pp* scattering data below $T_{\rm lab}$ = 350 MeV. In the description of all partial waves we take exactly into account the long-range potential consisting of the improved Coulomb potential (including the magnetic moment interaction), the vacuum polarization potential, and the tail of the one-pion-exchange potential. To describe the short-range interaction in the lower partial waves we use a $P$-matrix parametrization. The intermediate partial waves are treated either by optimal mapping techniques or by using the Nijmegen soft-core potential. The latter gives a better description of the data. The final data set comprises 1626 scattering observables. The best fit to this final data set results in $\chi^2/N_{\rm df}$ = 1.117, where $N_{\rm df}$ = 1576 is the number of degrees of freedom. The $pp\pi^0$ pseudovector coupling constant is determined to be $f_0^2 = (74.9 \pm 0.7) \times 10^{-3}$. Single-energy phase shifts and errors are also given.

PACS number(s): 21.30.+y, 13.75.Cs, 11.80.Et, 12.40.Qq



*Published in Phys. Rev. C **41**, 1435 (1990).

## I. INTRODUCTION

The phase-shift analysis of all $pp$ scattering data below $T_{\rm lab} = 350$ MeV presented here, is a continuation to higher energies of our 0–30 MeV analysis [1]. The $P$-matrix method that was there seen to be very successful in the description of the phase shifts as a function of the energy, is applied here also. The first results of the analysis have already been reported elsewhere [2, 3]. In this paper we want to give a full account of the methods and results of the analysis.

From the large, but still finite, number of scattering data one cannot determine an infinite number of phase shifts. However, the long-range part of the interaction is well known, whereas the unknown part of the interaction is sufficiently short ranged as to be screened by the centrifugal barrier. Therefore, all phase parameters with high orbital angular momentum $l$ are well known and we are left with only a finite number of phase shifts to be determined.

With the well-known long-range interaction we mean those effects that are theoretically well understood and that are the same in all realistic descriptions of the $pp$ interaction. We regard as such the tail of the electromagnetic potential, containing the relativistically corrected Coulomb potential [4] (including the magnetic moment interaction), the vacuum polarization potential [5], and the tail of the one-pion-exchange (OPE) potential. Not all phase-shift analyses include vacuum polarization [6] or Coulomb distortion in the higher partial waves [6, 7].

To give a non-OPE contribution to the partial waves with intermediate angular momentum $J$, we tried several methods [3]. We used optimal mapping techniques and we also tried to compute them directly via the Nijmegen soft-core (NSC) potential [8]. In all cases an improvement of the description of the data was obtained. The largest improvement was obtained with the NSC potential, which was therefore used in our final fits.

The main difference between this and other multienergy (ME) phase-shift analyses is, that we fully exploit the well-known long-range $pp$ interaction also in the description of the energy dependence of the phase parameters of the *lower* partial waves ($J \leq 4$). This is implemented by using outside some radius $r = b$ a potential tail in the relativistic radial Schrödinger equation. The phenomenology, necessary to describe the precise scattering data accurately, enters the method via the parametrization of a boundary condition at $r = b$ for which we use the $P$ matrix [9], which is the logarithmic derivative of the radial wave function at $r = b$. In our analysis, phenomenology is therefore only used where there is really a lack of knowledge, i.e., for $r < b$. This method has also an advantage when the long-range interaction is not fully determined and still contains some unknown parameters. These unknown parameters can then be determined from all partial waves. A good example is the OPE potential. The $pp\pi^0$ coupling constant $f_0^2$ has been determined this way [2] from a preliminary analysis of all $pp$ scattering data below 350 MeV.

In the treatment of the lower partial waves it appeared [3] that with OPE as the only nuclear interaction outside $r = b$, we had to take $b \geq 1.8$ fm in order to get a reasonable description of the data. With intermediate-range forces included in the potential tail, one can take a smaller value for $b$, the description of the data is better, and the parametrization of the $P$ matrix requires less parameters [3]. Although the description of the intermediate-range part of the nuclear interaction due to multiple-meson exchange or heavier-boson exchange (HBE) is model dependent, different potential models fortunately do agree very well for

$r \geq 1.4$ fm. This remarkable feature allows the use of such a potential for $r \geq 1.4$ fm without introducing too much model dependence. Since the tail of the Nijmegen soft-core (NSC) potential model [8] did give a slightly better fit than the tail of the parametrized Paris potential [10], we chose to include the HBE forces of the NSC model. With this model for the intermediate-range interaction, the results for the $P$ matrix, which parametrizes the interaction in the inner region, were not satisfactory for the $^1D_2$ and $^1G_4$ partial waves. Therefore, in order to allow for an adjustment of the used HBE potential model to remedy possible imperfections of the intermediate-range forces, the HBE forces for the singlet partial waves were multiplied with an arbitrary parameter, to be determined by the data. For the triplet partial waves such a parameter was not necessary. It should be emphasized that our method for analyzing the scattering data is especially suited to measure the quality of potential tails. In the future more nuclear potential tails will have to be tested in this way.

Our final multienergy fit to the data has $\chi^2/N_{\rm df} = 1.117$. This is considerably lower than other ME phase-shift analyses [6, 7] that have $\chi^2/N_{\rm df} \approx 1.3$. Therefore, we think that our ME phase parameters are more in accordance with the data. We also give single-energy (SE) phase shifts and errors. The full error matrices, which take into account the correlation between the phase parameters, and which should be used when adjusting the parameters of a model, are available upon request.

The outline of the paper is as follows. In Sec. II we discuss the method of analysis, based on the parametrization of the $P$ matrix, which is a boundary condition at $r = b$. Some properties of the $P$ matrix are briefly discussed in Sec. III and the treatment of the well-known long-range interaction ($r > b$) is presented in Sec. IV. Section V describes the data set we used. The determination of the neutral pion-nucleon coupling constant is discussed in Sec. VI, where we also present our multienergy and single-energy results and compare them with other phase-shift analyses.

## II. THE METHOD OF ANALYSIS

In previous multienergy (ME) phase-shift analyses various ways of parametrizing the phase parameters as a function of the energy have been employed. The simplest way is used by Bystricky, Lechanoine-Leluc, and Lehar [6]. They express the phase parameters as a polynomial of the energy, multiplied by the OPE phase parameters to ensure the correct threshold behavior. For the $^1S_0$ an exception is made; there the effective range approximation is used instead of the OPE phase shift. Arndt, Hyslop, and Roper [7] express the partial-wave amplitude as a sum over one-boson exchange type basis functions, where the masses are chosen as a multiple of the $\pi^0$ mass. The strengths of these amplitudes are then fitted to the data. Phase-shift analyses primarily interested in the low-energy region usually use effective range parametrizations or potential parametrizations. The drawbacks of these latter methods are extensively discussed in Ref. [1].

Our method of analysis is about the same as in our 0–30 MeV analysis [1]. What is special about this method? It is that we tried to exploit as much as possible our knowledge about the *pp* interaction. For large distances the electromagnetic interaction and the OPE part of the nuclear interaction are well known and model independent. Concerning the intermediate-range part of the nuclear interaction (say for $r \gtrsim 1$ fm), there is a remarkable

agreement between different potential models, despite the fact that there are significant differences in the heavier meson-nucleon coupling constants used, and in the treatment of the two-pion exchange. The short-range part of the nuclear interaction ($r \lesssim 1$ fm) is to a large extent unknown, as can also be seen in the differences between the different potential models.

Therefore, we view the *pp* interaction as being built up from several parts, which become more well known with larger $r$. The connection between the unknown inner region and the well-known outer region is implemented via a boundary condition model. The long-range and intermediate-range interactions are incorporated via a potential tail outside $r = b$ in the relativistic radial Schrödinger equation [11], which is nothing else but a coordinate space version of the relativistic Lippmann-Schwinger integral equation, which in turn is totally equivalent with three-dimensional integral equations, such as the Blankenbecler-Sugar equation [12]. The unknown short-range interaction is parametrized phenomenologically by specifying the $P$ matrix at $r = b$ in each partial wave. The long-range electromagnetic and the OPE potential tails are taken into account exactly. Different models for the intermediate-range forces have been tried. In the final analysis a specific model for these forces has been chosen.

The asymptotic behavior of the radial wave function, which is the solution of the radial Schrödinger equation with specified boundary condition at $r = b$, gives the partial wave $S$ matrix by matching the wave function to Coulomb functions. The $S$ matrix is decomposed into the standard nuclear-bar phase shifts and mixing parameters [13]. For the definition of different kinds of phase parameters, found by matching the radial wave function to functions other than Coulomb functions, see Ref. [1].

For larger angular momentum the inner region becomes less important due to the centrifugal barrier. Therefore, in a phase-shift analysis, one usually parametrizes the lower partial waves only, whereas the higher partial-wave phase parameters are approximated by the OPE phase parameters. The quality of the analysis can be improved significantly if one has a better approximation than OPE for the partial waves with intermediate angular momentum. In a preliminary analysis [3], we have used the optimal polynomial theory (OPT) [14, 15] to give a non-OPE contribution to the phase parameters. There, OPT is only used for the ${}^1S_0$, ${}^3P_1$, and ${}^3P_2$ sequences of partial waves, since in our study of OPT [3], we found that it could neither predict the mixing parameters nor the ${}^3F_2$ sequence well enough. Another possibility is to use the heavier boson exchange (HBE) forces of some potential model. In our present analysis we used the HBE forces of the Nijmegen soft-core (NSC) model to give a non-OPE contribution to the phase parameters with intermediate angular momentum.

Given all partial-wave phase parameters, the observables are computed in a standard way [16, 17]. The difference of the observables with the experimental data is minimized in a $\chi^2$ fit. Some data are rejected on the basis of statistical criteria. The number of parameters is determined by the criterion that the fit does not improve significantly when a parameter is added.

In any ME analysis problems can be expected at the end of the energy range. In our analysis, such a problem was most strikingly seen for the ${}^1S_0$. The ${}^1S_0$ phase shift showed a tendency to rise at the highest energies. Since it is well known from other phase-shift analyses [6, 7, 18, 19, 20], that the ${}^1S_0$ phase shift continues to drop smoothly as a function

of energy up to energies far above 350 MeV, we added the $^1S_0$ phase shift at 425 MeV as an extra datum. The value chosen, $\delta(^1S_0) = -19 \pm 2^\circ$ at 425 MeV, has a sufficiently large error such that it is in agreement with all the aforementioned higher-energy analyses but still insures a proper energy dependence near 350 MeV.

Adjusting all parameters to obtain the $\chi^2_{\rm min}$ value gives our final ME fit. The quality of the fit can be seen from the $\chi^2_{\rm min}$ value, and from the error matrix of the parameters, which is the inverse of half the second derivative matrix of $\chi^2_{\rm min}$ with respect to the parameters.

The energy dependence of the phase parameters in the final ME fit is used in the single-energy (SE) analyses. For the SE analyses we divided the data into ten clusters, from which the SE phase parameters and inverse error matrices were determined. For each phase parameter searched for, we fitted a constant to be added to the energy-dependent $P$ matrix of the ME fit. The results of the SE analyses, i.e., SE phase parameters and error matrices, are a compact representation of the $\chi^2_{\rm min}$ surface. They can be used by anyone who wants to adjust the parameters of a model for the $pp$ interaction to the data. Of course, modelists can also compare directly with the data, but this has its disadvantages [21]. Different modelists will pick different sets of data, sometimes perhaps only data that give a low $\chi^2$ value. Comparison of the quality of these models will then be troublesome.

## III. THE P MATRIX

### A. Definition and properties of the P matrix

We briefly review the definition and properties of the $P$ matrix. For a more detailed discussion, see Ref. [1].

For $r > b$ we use the relativistic [11] radial Schrödinger equation

$$\left[\frac{d^2}{dr^2} + k^2 - \frac{L^2}{r^2} - M_p V(r)\right] \chi(r) = 0 \ , \tag{1}$$

where $\chi(r)$ is the radial wave function, $M_p$ is the proton mass, and $L^2$ is a shorthand notation for $l(l+1)$, with $l$ the orbital angular momentum. The correct relativistic relation between the center-of-mass (c.m.) relative momentum $k$ and the laboratory kinetic energy $T_{\rm lab}$ is $k^2 = M_p T_{\rm lab}/2$.

The interaction inside $r = b$ is described by a boundary condition at $r = b$, the $P$ matrix

$$P(b; k^2) = b \left[\frac{d\chi}{dr}\chi^{-1}\right]_{r=b} \ . \tag{2}$$

Given the asymptotic behavior of $\chi(r)$, and the potential $V(r)$ outside $r = b$, the $P$ matrix is uniquely determined.

We add the well-known long-range interaction by means of a potential tail and parametrize the structure of the $P$ matrix as a function of the energy. The potential $V(r)$ we use for $r \geq b$ is discussed in Sec. IV, and the parametrizations for the $P$ matrix are discussed in Sec. III B.

The property of the $P$ matrix on which the parametrizations in this analysis are based is that, if one assumes that a local potential $V(r)$ exists for $r < b$ also, the $P$ matrix can be written as a sum of poles. In the one-channel case we may write

$$P(b;k^2) = c + k^2 \sum_{n=1}^{\infty} \frac{r_n}{k^2 - k_n^2} \ . \tag{3}$$

For comparison, one might look at the trivial case that $V(r) = 0$ for $r < b$. For the partial wave with orbital angular momentum $l$ this leads to what we call the free $P$ matrix $P_{\rm free}$ for which

$$c = l + 1, \quad r_n = 2, \quad k_n = z_n/b \ , \tag{4}$$

with $z_n$ the $n$th zero of the spherical Bessel function $j_l(z)$.

In the case that Eq. 1 contains a potential which has a constant value $V$ for $r < b$, this can be absorbed in the $k^2$ term and we then have

$$P(b;k^2) = P_{\rm free}(b;k^2 - M_p V) \ . \tag{5}$$

For a coupled-channels $P$ matrix, the residues of Eq. (3) can be factorized. This means that in the neighborhood of a pole one may write

$$P_{ij} = (P_{\rm bg})_{ij} + \xi_i \xi_j \frac{1}{k^2 - k_n^2} \ , \tag{6}$$

where $P_{\rm bg}$ describes a smooth background. This means that in general the pole will show up in all matrix elements, with residues that are not independent. This is of importance for the choice of the $P$-matrix parametrization.

The potential that we use will not be entirely exact, since the short-range and intermediate-range forces are not well known and model dependent. Furthermore, our potential tail does not include inelasticities. Therefore, we cannot expect the above properties to hold exactly. The $S$ matrix has a unitarity cut, right-hand cuts due to inelasticities, and left-hand cuts due to particle exchanges. In the $P$ matrix, on the other hand, some of these cuts are removed by including the correct potential tail. Therefore, the $P$ matrix used in our analysis has its nearest left-hand cut at $T_{\rm lab} = -38.83$ MeV. This cut is due to those two-pion exchange effects that are not included in the iterated one-pion exchange or heavier-boson exchange potentials. The lowest-lying right-hand cuts due to the inelasticities and the $T_{\rm lab}$ in MeV of the corresponding thresholds are $pp\pi^0$ (279.6), $d\pi^+$ (287.5), and $pn\pi^+$ (292.3). One might expect to find some effects of these thresholds at the high end of our energy range. At energies below 400 MeV, however, the inelasticity appears to be still rather small [6, 7, 18], so we will neglect it. An improvement of our model would be the inclusion of the right-hand cuts in some way. This will be necessary if higher energies are to be included.

### B. Parametrizations of the P matrix

Many parametrizations of the $P$ matrix can be used to obtain a phenomenological description of the short-range interaction. We start by looking at an uncoupled channel. One

could use Eq. (3) and parametrize the $P$ matrix by a finite number of poles. However, this is not the best method if the energy range that is analyzed contains no pole, since in that case a distant pole outside the energy range does not affect the $P$ matrix very much, making it impossible to fit the two parameters $r_n$ and $k_n$ of Eq. (3) to the data. The effect of all the higher poles together can be taken into account by adding a power series in $k^2$. This leads to

$$P = \frac{k^2 r}{k^2 - k_p^2} + \sum_{n=0}^{N} c_n (k^2)^n \ . \tag{7}$$

A simpler parametrization can be obtained by using no pole at all and keeping only the second term in Eq. (7).

A different parametrization is obtained by starting with $P_{\text{free}}$, the free $P$ matrix, and replacing the argument $k^2$ by some function $f(k^2)$. If $f$ is expanded in $k^2$ this leads to the parametrization

$$P = P_{\text{free}} \left[ b; k^2 - \sum_{n=0}^{N} a_n (k^2)^n \right] \ . \tag{8}$$

Using Eq. (5) for $r < b$, we see that we have in fact used an energy-dependent potential, independent of $r$,

$$V(k^2) = \frac{1}{M_p} \sum_{n=0}^{N} a_n (k^2)^n \ . \tag{9}$$

This is the parametrization we use in our final ME fit. Instead of the $k^2$-dependent but $r$-independent potential of Eq. (9), one can of course use any short-range potential to parametrize the $P$ matrix.

In the $pp$ interaction we encounter the case of two coupled channels. To describe this coupling, as far as the short-range interaction is involved, we need to construct a $2 \times 2$ $P$ matrix. The simplest way to do so starts with two single-channel $P$ matrices as diagonal elements. These can be constructed using one of the methods mentioned earlier. After that one can simply add the off-diagonal elements as some function of $k^2$, e.g., using a power series

$$P = \begin{bmatrix} P_1 & \sum_{n=0}^{N} d_n (k^2)^n \\ \sum_{n=0}^{N} d_n (k^2)^n & P_2 \end{bmatrix} \ . \tag{10}$$

The $P$ matrix constructed in this way will not have the correct behavior in the vicinity of a pole [see Eq. (6)]. A better construction is

$$P = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} P_1 & 0 \\ 0 & P_2 \end{pmatrix} \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \ , \tag{11}$$

where the angle $\theta$ is a smooth function of $k^2$, which can be expanded in a power series. Equation (11) must be used if poles appear in the analyzed energy range. It can also be expected to be better than Eq. (10) if the poles are close to the energy range.

## IV. THE POTENTIAL TAIL

The long-range interaction, beyond some distance $b$, is described by a potential tail. Our objective is to include at least all potentials that can be considered to be model independent.

The electromagnetic interaction is almost the same as in our 0–30 MeV phase-shift analysis [1]; next to the relativistically corrected Coulomb potential $\tilde{V}_{\rm C}$ [4] and the vacuum polarization potential $V_{\rm VP}$ [5], we here have also included the magnetic-moment interaction $V_{\rm MM}$ of the improved Coulomb potential [4, 22]. For $T_{\rm lab} > 30$ MeV vacuum polarization is unimportant. Explicitly, the included electromagnetic potential $V_{\rm EM}$ is

$$V_{\rm EM} = \tilde{V}_{\rm C} + V_{\rm MM} + V_{\rm VP} = V_{\rm C1} + V_{\rm C2} + V_{\rm MM} + V_{\rm VP} \ , \tag{12}$$

with

$$\begin{aligned} V_{\rm C1} &= \frac{\alpha'}{r} \ , \\ V_{\rm C2} &= -\frac{1}{2M_p^2}\left[(\Delta + k^2)\frac{\alpha}{r} + \frac{\alpha}{r}(\Delta + k^2)\right] \\ V_{\rm MM} &\sim -\frac{\alpha}{4M_p^2 r^3}\left[\mu_p^2 S_{12} + (6 + 8\kappa_p)\mathbf{L}\cdot\mathbf{S}\right] \ , \end{aligned} \tag{13}$$

where $\alpha' = 2k\eta'/M_p$ with $\eta' = \alpha/v$lab the standard Coulomb parameter [23], $\Delta$ is the Laplacian, $\mu_p = 2.792\,847$ is the proton magnetic moment, and $\kappa_p = \mu_p - 1$ is the anomalous magnetic moment.

The magnetic-moment potential is obtained from the phenomenological interaction Lagrangian

$$\mathcal{L}_V = e\overline{\psi}\left(F_1 i\gamma^\mu A_\mu + F_2\sigma^{\mu\nu}\partial_\mu A_\nu\right)\psi \ , \tag{14}$$

where $\sigma^{\mu\nu} = [\gamma^\mu, \gamma^\nu]/2i$, $A_\mu$ is the vector-meson (photon) field, and $\psi$ is the proton field. The Dirac and Pauli form factors are given by

$$F_1 = \frac{1 + \mu_p\tau}{1 + \tau}G_E \ , \quad F_2 = \frac{\kappa_p}{2M_p}\frac{G_E}{1 + \tau} \ , \tag{15}$$

where $\tau = -t/4M_p^2$ with $t$ the Mandelstam momentum transfer, and $G_E$ is the Sachs proton electric form factor. Using well-known techniques [11] the one-photon-exchange momentum space $pp$ potential is derived. Introducing the standard momentum dependence for the Sachs form factor as obtained from the dipole fit by Hofstadter and co-workers [24]

$$G_E(t) = \frac{1}{(1 - t/m_D^2)^2} \ ,$$

with $m_D^2 = 0.71(\mathrm{GeV}/c)^2$ the "dipole" mass squared, the potential is Fourier transformed to configuration space. The momentum dependence of the form factors gives rise to short-range Yukawa-like central, spin-spin, tensor, and spin-orbit potentials. The long-range part of the magnetic-moment potential is given as in Eq. (13). The contribution of $V_{\rm MM}$ to the scattering amplitude is obtained by summing the potential, integrated with Coulomb

functions, for each partial wave up to $l = 1000$, since one should use the Coulomb distorted-wave Born approximation (CDWBA) rather than the plane wave Born approximation (BA) in calculating the $pp$ magnetic-moment scattering amplitude. This was already pointed out by Knutson and Chiang [25]. The spin-orbit part of the interaction, however, gives rise to a term that converges too slowly for a summation on a computer to be practical. Fortunately, this part can be summed analytically (see also Ref. [25]). A more detailed account of our treatment of the magnetic-moment interaction in our phase-shift analysis will be published elsewhere [22].

The longest-range nuclear potential, one-pion exchange, is included as

$$V_{\rm OPE} = \frac{1}{3} f_0^2 \, (\frac{m}{m_c})^2 \, \frac{M_p}{E} \frac{e^{-mr}}{r} \times \left[ (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) + S_{12} \left( 1 + \frac{3}{(mr)} + \frac{3}{(mr)^2} \right) \right] , \tag{16}$$

where $m$ is the $\pi^0$ mass, $m_c$ is a scaling mass, conventionally chosen to be the charged pion mass [26], and $E = (M_p^2 + k^2)^{1/2}$ with $k$ the c.m. relative momentum. No form factor has been included, since the latter represents a short-range effect, whereas the potential is used for $r > b$ only. The $pp\pi^0$ pseudovector coupling constant $f_0^2 \equiv f_{pp\pi^0}^2/4\pi$, appearing in this potential, is left as a free parameter. It determines the strength of the Yukawa tail, so it represents the coupling constant at the pion pole. The contributions of the long-range potentials $V_{\rm EM}$ and $V_{\rm OPE}$ to the scattering amplitude are calculated in CDWBA.

The intermediate-range forces cannot be included without introducing some model dependence. Therefore, we first did not include any nuclear potential tail other than OPE. A reasonable fit was then only possible for $b \geq 1.8$ fm. When we [3] used the intermediate-range forces of the Nijmegen soft-core potential [8] or the parametrized Paris potential [10], $b$ could be chosen smaller and the fit to the data became better, even if less parameters were used for the short-range interaction. These potentials were used in our model after removing their OPE part, keeping only what we call "intermediate-range forces." In all cases we kept the OPE potential of Eq. (16). Since the fit to the data with the Nijmegen intermediate-range forces was somewhat better than with the Paris intermediate-range forces [2], we chose to include the non-OPE part of the Nijmegen potential in our potential tail for our final ME fit.

Although an improved fit was obtained in this way, there were still indications that this potential tail was not perfect. This was most clearly seen in the ${}^1D_2$ and ${}^1G_4$ partial waves. The potential in the inner region, which is the equivalent of the $P$ matrix, appeared to be highly attractive for these waves. This points towards an incorrect potential tail. This is because of the fact that the region $r < b$ has only little influence on the phase shifts for higher $l$, so in order to compensate for rather small imperfections in the potential tail, very large short-range potentials are needed. To investigate this we multiplied the intermediate-range potential by the factor $f_{\rm med}$, which allows for an adjustment of the used potential model to remedy possible imperfections of the intermediate-range forces. This parameter $f_{\rm med}$ makes the tail of the potential more attractive, so one needs less exorbitant attraction in the inner region. The non-OPE part of the Nijmegen potential ($V_{\rm HBE}^N$) is multiplied for the singlet waves with this factor, now called $f_{\rm med}^s$. The best fit was achieved with $f_{\rm med}^s \approx 1.6$. For the triplet waves such a parameter turned out to be unnecessary (i.e., $f_{\rm med}^t \approx 1.0$).

Summarizing, the employed potential tail $V_L$ is

$$V_L = V_{\rm EM} + V_{\rm NUC} = V_{\rm EM} + V_{\rm OPE} + V^N_{\rm HBE}(f^s_{\rm med}) \ , \tag{17}$$

and contains two parameters, $f_0^2$ and $f^s_{\rm med}$. In our final fit $f^s_{\rm med}$ was fixed at 1.6 to save computer storage and time, whereas $f_0^2$ was still left as a free parameter.

## V. THE DATA

A list of all groups of data we used in our analysis is shown in Table I. Although most of these data already have been presented in previous analyses [6, 18, 19, 128, 129, 130, 131] we give them here explicitly, since we wish to provide a self-contained and complete data base for our future phase-shift analyses.

Our data set consists of all *pp* scattering data below $T_{\rm lab} = 350$ MeV, published in a regular physics journal as of 1955 (we therefore rather arbitrarily do no longer include the Princeton54 cross-section data of Yntema and White [132], which were published in 1954, although they were included in our 0–30 MeV analysis [1]), and is updated up to August 1989. A detailed list of the major part of the data can be found in the Nucleon-Nucleon Scattering Data Tables of Bystricky and Lehar [127, 122] as published in 1978 and 1981. For all experimental results we have consulted the original references and we corrected for some minor printing errors in the Scattering Data Tables. Moreover, we include data groups that are not contained in these tables (e.g., because they were published after 1981). These groups are denoted by a dagger. Groups of data that are not included in the latest data set NN896 of SAID (Ref. [123]) are denoted by an asterisk.

On the other hand, we do not include dispersion relation predictions [133] and data that were obtained from quasielastic scattering (e.g., deuteron targets). Total cross-section data $(\sigma_{\rm tot}, \Delta\sigma_T, \Delta\sigma_L)$ were omitted, because of the difficulties of their definition and because of the differences in the treatment of the Coulomb-nuclear interference term by the experimentalists. As in our 0–30 MeV analysis [1], we also do not include data that have not been published (yet) or that have only been reported in conference proceedings. The main difference between our data base and that of other analyses [6, 7, 18] is that we have extended the energy range downwards to 0 MeV and included all low-energy data. The data set up to 30 MeV is almost the same as in our 0–30 MeV analysis [1]. However, we added one polarization measurement [44], one spin correlation [48], and 22 cross-section data [46, 47], which were found to be missing from our previously published analysis of the *pp* scattering data below $T_{\rm lab} = 30$ MeV.

Starting with this set of 1917 scattering observables, the $P$-matrix parameters and the neutral pion-nucleon coupling constant were adjusted to obtain a $\chi^2_{\rm min}$. Data that were more than 3 standard deviations ($3\sigma$) off were rejected and the parameters were adjusted again. If the experimental normalization on a group of data contributed more than 9 to $\chi^2$, these data were floated (freely renormalized) by us, which is indicated by "$n$" in the column labeled "Comment" in Table I. The original experimental normalization error is shown in parentheses for such cases. Groups which had an improbably low or high $\chi^2$ were rejected also. Groups rejected this way are indicated by "$j$" and "$o$," respectively, in the column labeled "Comment."

If we compare with our 0–30 MeV analysis [1], we note that the Berkeley67 polarization data of Slobodrian *et al.* [37] are no longer rejected, except for the six polarization data at 19.7 MeV that still have a $\chi^2$ that is too high. All other changes in the data set below $T_{\rm lab} = 30$ MeV result in the same data set as used in our 0–30 MeV analysis.

Making all these adjustments to the complete data set results in a total number of 291 data to be rejected, leaving us with our final data set of 1626 scattering observables divided over 213 groups, as presented in Table I. Of these groups 117 have an experimental normalization error and 22 groups have a floated normalization. Each group of Zürich78 cross-section data [28] contains two angle-dependent normalization data. Including the 27 $P$-matrix parameters and the pion-nucleon coupling constant, we are left with a total number of degrees of freedom of $N_{\rm df} = 1576$. The parameter $f^s_{\rm med}$ and the boundary condition radius were fixed at $f^s_{\rm med} = 1.6$ and $b = 1.4$ fm in our final ME fit and we do not count them as free parameters.

## VI. RESULTS

Starting with the complete dataset of 1917 $pp$ scattering observables, the 27 $P$-matrix parameters, the pion-nucleon coupling constant $f_0^2$, the free parameter $f_{\rm med}$, indicating the strength of the HBE potential tail outside $r = b$, were adjusted to obtain the best possible description of the data. The boundary condition radius is set at $b = 1.4$ fm. Of the data, 291 did not survive our rejection criteria, so our final data set contains 1626 scattering observables (with an additional 129 normalization data). The parameter $f_{\rm med}$ was found to be helpful in the singlet partial waves only and was fixed in our final ME fit to be $f^s_{\rm med} = 1.6$.

Taking into account all free parameters, floated normalizations and other normalization errors, we are left with a total number of degrees of freedom of $N_{\rm df} = 1576$. Our minimum $\chi^2$ value is $\chi^2_{\rm min} = 1760.6$ if we include the magnetic-moment interaction and $\chi^2_{\rm min} = 1789.2$ if we do not include it, which means a five standard deviation effect. We do not find a significant difference in $\chi^2_{\rm min}$ if we replace the momentum dependent (dipole) proton form factors by their point particle approximation, which is not surprising in view of the short-range effect of this momentum dependence. We obtain $\chi^2/N_{\rm df} = 1.117$ for the analysis with the magnetic-moment interaction included, and $\chi^2/N_{\rm dat} = 1.003$, where $N_{\rm dat} = 1755$ is the number of data points (scattering observables and normalization errors). This is lower than any other phase-shift analysis [6, 7, 18, 19, 20].

The $pp\pi^0$ pseudovector coupling constant is found to be

$$f_0^2 = (74.9 \pm 0.7) \times 10^{-3} \quad \text{or} \quad g_0^2 = 13.55 \pm 0.13 \ ,$$

where

$$g_0^2 = (2M_p/m_c)^2 f_0^2 = 180.78 f_0^2 \ .$$

This result is slightly higher than our previously published value of $f_0^2 = (72.5 \pm 0.6) \times 10^{-3}$, or $g_0^2 = 13.1 \pm 0.1$, obtained from our preliminary earlier phase-shift analysis [2]. However, that analysis did not contain the magnetic-moment interaction. Because of its long range, the magnetic-moment interaction is likely to influence the value of $f_0^2$ so we have to compare

the result with the result from the analysis without the magnetic-moment interaction. This analysis now yields

$$f_0^2 = (73.7 \pm 0.7) \times 10^{-3} \ ,$$

or $g_0^2 = 13.33 \pm 0.13$, in good agreement with our preliminary analysis [2].

The result for $f_0^2$ is in reasonable agreement with earlier (less accurate) determinations (see Table II), except for the value obtained by Kroll [133]. It differs, however, by more than three standard deviations from the value for the charged coupling constant as determined from $\pi N$ scattering [26]

$$f_c^2 = (78.9 \pm 1.0) \times 10^{-3}$$

or $g_c^2 = 14.28 \pm 0.18$, where

$$g_c^2 = [(M_p + M_n)/m_c]^2 f_c^2 = 181.03 f_c^2 \ .$$

For charge independence [SU(2)-isospin symmetry] of the pion-nucleon interaction one has $f_0^2 = f_c^2$, so our determination of the $pp\pi^0$ coupling constant seems to indicate a large breaking of charge independence.

Introducing different coupling constants for the spin triplets $f_T$ and for the spin singlets $f_S$, we obtain

$$f_S^2 = (76.0 \pm 1.3) \times 10^{-3}$$

and

$$f_T^2 = (74.8 \pm 0.7) \times 10^{-3} \ ,$$

indicating the importance of the spin-triplet waves in the determination. When we next introduce different coupling constants for the ${}^3P$ waves $f({}^3P)$ and all other partial waves $f(\text{rest})$, we find

$$f({}^3P)^2 = (74.7 \pm 0.8) \times 10^{-3}$$

and

$$f(\text{rest})^2 = (75.1 \pm 0.7) \times 10^{-3} \ ,$$

indicating that the ${}^3P$ waves are not especially important in the determination of the coupling constant. It is therefore not possible for us to pinpoint some special type of observables as being responsible for the low value of $f_0^2$. It rather appears that the data as a whole require a low pion-nucleon coupling constant.

The multienergy results for each group of data is given in Table I. There we give the $\chi^2$-values, the predicted normalization with which the experimental data should be multiplied before comparing them with the theoretical values, which data were rejected, and why they were rejected.

The phase parameters were parametrized with 27 $P$-matrix parameters. For the $P$-matrix parametrization we used an energy-dependent potential, independent of $r$ for $r < b$ [Eqs. (8) and (11)]. We allow for up to seven parameters in each partial wave (i.e., up to sixth order in $k^2$), which was found to be enough (only the $^1S_0$ required all seven parameters). The resulting coefficients as obtained in our final ME fit are presented in Table III. The coefficients $c_n$ have dimension [fm$^{2n}$] and are related to the coefficients of Eq. (9) according to $a_n = c_n/b^2$, i.e., the corresponding energy-dependent potential in MeV is obtained by multiplying with $1/M_p b^2 = 21.173$ MeV. The ME phase parameters were used in the computer code SAID (Ref. [123]) to provide an objective criterion for the quality of our results. For 1083 data in the energy range 3–325 MeV this results in $\chi^2 = 1092.2$ or $\chi^2/N_{\rm dat} = 1.01$, in excellent agreement with the result of our own phase-shift analysis of $\chi^2/N_{\rm dat} = 1.02$ in the same energy range.

Single-energy phase-shift analyses were performed by fitting a constant to the $P$-matrix parametrization (which was now held fixed at the ME result in order to insure the proper local energy behavior) for all phase parameters that were to be determined. The SE analyses were performed at ten different energies from 0.382 54 MeV (the interference minimum) up to 320 MeV. The data were divided into ten groups and clustered around the appropriate energies. The results are presented in Table IV, where we give the ME $\chi^2$, the ME phase parameters, the SE $\chi^2$, and the SE phase parameters with their errors. The errors are the square roots of the diagonal elements of the inverse of one half times the second derivative matrix of $\chi^2$ with respect to the parameters (the error matrix). Actually, all phase parameters within a SE analysis are correlated with each other, so instead of the error as presented in Table IV, one should use the complete error matrix whenever one wants to determine quantities derived from a combination of phase parameters. Since these error matrices become larger with the number of searched phase parameters ($12 \times 12$ matrices for the SE analyses at 215 and 320 MeV) we do not reproduce them here because of lack of space. They are available upon request. The agreement between the ME and SE phase parameters is good, which validates our procedure of using the energy slopes from the ME analysis when doing the single-energy analysis.

The phase parameters are phase shifts and mixing parameters with respect to electromagnetic (EM) wave functions ($\delta^{C+EM}_{C+EM+N}$ in the notation of Ref. [1]). They can be converted to phase shifts with respect to Coulomb functions by subtracting the EM phase shifts $\delta^C_{C+EM}$, which contain the contributions of the relativistic Coulomb interaction, the magnetic-moment interaction with dipole form factors, and the vacuum polarization. Since the magnetic-moment interaction contains a tensor part [see Eq. (13)] one should use the $S$ matrix whenever one wants to convert the triplet coupled phase parameters with respect to EM wave functions to phase parameters with respect to Coulomb functions, i.e.,

$$S^C_{C+EM+N} = (S^C_{C+EM})^{1/2} S^{C+EM}_{C+EM+N} (S^C_{C+EM})^{1/2} \ .$$

Doing this results in the corrections $\delta^C_{C+EM}$ for each phase parameter separately as given in Table IV. The phase parameters can then be compared with the results of our 0–30 MeV analysis [1]. The differences remain well within one standard deviation for most phase parameters.

Table IV furthermore contains the number of scattering observables and the number of degrees of freedom for each SE analysis.

The ME phase parameters are shown in Fig. 1 together with the results of the phase-shift analyses of Arndt, Hyslop, and Roper [7], Bystricky, Lechanoine-Leluc, and Lehar [6], Dubois *et al.* [18], and Bugg *et al.* [19]. All phase parameters can be determined rather accurately and compare favorably with the other phase-shift analyses. However, we think that our phase parameters are more in accordance with the data, because of our considerably lower value of $\chi^2/N_{\rm df}$ as compared to, e.g., Arndt [7] ($\chi^2/N_{\rm df} = 1.3$). Furthermore, the bulk of new data of Onel *et al.* [103] at 241, 314, and 341 MeV which have not been used in previous phase-shift analyses, provide important constraints on the phase parameters at higher energies. Especially the ${}^3F$ phase shifts above 200 MeV, that can now be determined more accurately than in previous analyses. The differences with, e.g., Arndt are most strikingly seen in the ${}^3F_2$ and ${}^3F_3$ phase shifts at 300 MeV.

The central, tensor, and spin-orbit combinations of the $P$ and $F$ waves are shown in Fig. 2. The SE results are given with their error bars, whereas the other analyses do not publish error bars for their combined phase parameters. The spin-orbit combinations of both $P$ and $F$ waves are determined very accurately.

## VII. SUMMARY

Summarizing, we have performed a phase-shift analysis of all $pp$ scattering data below $T_{\rm lab} = 350$ MeV, published in a regular physics journal as of 1955. The total data set contains 1917 scattering observables, of which 291 do not survive our rejection criteria. The final data set contains 1626 scattering observables and 129 normalization data. With 27 $P$-matrix parameters and the $pp\pi^0$ coupling constant we arrive at $\chi^2_{\rm min} = 1.117$, which is lower than any other phase-shift analysis.

Our method for analyzing the data allows us to determine the $pp\pi^0$ coupling constant at the pion pole and we find $f_0^2 = (74.9 \pm 0.7) \times 10^{-3}$. This result is significantly different from the charged pion-nucleon coupling constant $f_c^2 = (78.9 \pm 1.0) \times 10^{-3}$, indicating a large breaking of charge independence in the nucleon-nucleon interaction. More conclusive evidence for such a large breaking would be obtained if both neutral and charged coupling constants could be determined in a combined $pp$ and $np$ phase-shift analysis of all $NN$ scattering data below $T_{\rm lab} = 350$ MeV. This is under investigation.

The present results will serve as a basis for our future $NN$ phase-shift analyses. Some of the results of our $np$ phase-shift analysis below $T_{\rm lab} = 30$ MeV have already been published [136, 137], and the 0–350 MeV $np$ analysis is in progress. The next step will be to extend our analyses to energies above the threshold by including inelasticities.

## ACKNOWLEDGMENTS

Part of this work was included in the research program of the Stichting voor Fundamenteel Onderzoek der Materie (FOM) with financial support from the Nederlandse Organisatie voor Wetenschappelijk Onderzoek (NWO).

# REFERENCES


[1] J.R. Bergervoet, P.C. van Campen, W.A. van der Sanden, and J.J. de Swart, Phys. Rev. C **38**, 15 (1988).

[2] J.R. Bergervoet, P.C. van Campen, T.A. Rijken, and J.J. de Swart, Phys. Rev. Lett. **59**, 225 (1987).

[3] J.R. Bergervoet, P.C. van Campen, T.A. Rijken, and J.J. de Swart, Phys. Rev. C **38**, 770 (1988).

[4] G.J.M. Austen and J.J. de Swart, Phys. Rev. Lett. **50**, 2039 (1983).

[5] L. Durand III, Phys. Rev. **108**, 1597 (1957).

[6] J. Bystricky, C. Lechanoine-Leluc, and F. Lehar, J. Phys. (Paris) **48**, 199 (1987).

[7] R.A. Arndt, J.S. Hyslop III, and L.D. Roper, Phys. Rev. D **35**, 128 (1987).

[8] M.M. Nagels, T.A. Rijken, and J.J. de Swart, Phys. Rev. D **17**, 768 (1978).

[9] R.L. Jaffe and F.E. Low, Phys. Rev. D **19**, 2105 (1979).

[10] M. Lacombe, B. Loiseau, J.M. Richard, R. Vinh Mau, J. Côté, P. Pirès, and R. de Tourreil, Phys. Rev. C **21**, 861 (1980).

[11] M.H. Partovi and E.L. Lomon, Phys. Rev. D **2**, 1999 (1970); K. Erkelenz, Phys. Rep. **13C**, 191 (1974); J.J. de Swart and M.M. Nagels, Fortschr. Phys. **26**, 215 (1978).

[12] A.A. Logunov and A.N. Tavkhelidze, Nuovo Cimento **29**, 380 (1963); R. Blankenbecler and R. Sugar, Phys. Rev. **142**, 1051 (1966); V.G. Kadyshevsky, Nucl. Phys. **B6**, 125 (1968); Nuovo Cimento **55**, 275 (1968).

[13] H.P. Stapp, T.J. Ypsilantis, and N. Metropolis, Phys. Rev. **105**, 302 (1957).

[14] R.E. Cutkosky and B.B. Deo, Phys. Rev. **174**, 1859 (1968).

[15] S. Ciulli, Nuovo Cimento A **61**, 787 (1969); **62**, 301 (1969).

[16] M.H. MacGregor, M.J. Moravcsik, and H.P. Stapp, Annu. Rev. Nucl. Sci. **10**, 291 (1960).

[17] N. Hoshizaki, Prog. Theor. Phys. Suppl. **42**, 107 (1968).

[18] R. Dubois, D. Axen, R. Keeler, M. Comyn, G.A. Ludgate, J.R. Richardson, N.M. Stewart, A.S. Clough, D.V. Bugg, and J.A. Edgington, Nucl. Phys. **A377**, 554 (1982).

[19] D.V. Bugg, J.A. Edgington, C. Amsler, R.C. Brown, C.J. Oram, K. Shakarchi, N.M. Stewart, G.A. Ludgate, A.S. Clough, D. Axen, S. Jaccard, and J. Vávra, J. Phys. G **4**, 1025 (1978).

[20] R.A. Arndt, L.D. Roper, R.A. Bryan, R.B. Clark, B.J. VerWest, and P. Signell, Phys. Rev. D **28**, 97 (1983).

[21] J.J. de Swart, W.A. van der Sanden, and W. Derks, Nucl. Phys. **A416**, 299c (1984).

[22] V.G.J. Stoks and J.J. de Swart, Phys. Rev. C **42**, 1235 (1990).

[23] G. Breit, Phys. Rev. **99**, 1581 (1955).

[24] R. Hofstadter, F. Bumiller, and M.R. Yearian, Rev. Mod. Phys. **30**, 482 (1958).

[25] L.D. Knutson and D. Chiang, Phys. Rev. C **18**, 1958 (1978).

[26] O. Dumbrajs, R. Koch, H. Pilkuhn, G.C. Oades, H. Behrens, J.J. de Swart, and P. Kroll, Nucl. Phys. **B216**, 277 (1983).

[27] J.E. Brolley, Jr., J.D. Seagrave, and J.G. Beery, Phys. Rev. **135**, B1119 (1964); M.L. Gursky and L. Heller, *ibid.* **136**, B1693 (1964).

[28] Ch. Thomann, J.E. Benn, and S. Münch, Nucl. Phys. **A303**, 457 (1978).

[29] H. Wassmer and H. Mühry, Helv. Phys. Acta **46**, 626 (1973).

[30] J. Birchall, E. Baumgartner, H.M. Friess, H. Mühry, F. Rösel, and D. Trautmann, Helv. Phys. Acta **50**, 509 (1977).
[31] D.J. Knecht, P.F. Dahl, and S. Messelt, Phys. Rev. **148**, 1031 (1966).
[32] K. Imai, K. Nisimura, N. Tamura, and H. Sato, Nucl. Phys. **A246**, 76 (1975).
[33] M.D. Barker, P.C. Colby, W. Haeberli, and P. Signell, Phys. Rev. Lett. **48**, 918 (1982); **49**, 1056E (1982).
[34] R.J. Slobodrian, H.E. Conzett, E. Shield, and W.F. Tivol, Phys. Rev. **174**, 1122 (1968).
[35] G. Bittner and W. Kretschmer, Phys. Rev. Lett. **43**, 330 (1979).
[36] H. Obermeyer, H. Kuiper, B. Seidler, and K. Frank, Nucl. Phys. **A344**, 75 (1980).
[37] R.J. Slobodrian, J.S.C. McKee, H. Bichsel, and W.F. Tivol, Phys. Rev. Lett. **19**, 704 (1967).
[38] L.H. Johnston and D.E. Young, Phys. Rev. **116**, 989 (1959).
[39] N. Jarmie, R.E. Brown, R.L. Hutson, and J.L. Detch, Jr., Phys. Rev. Lett. **24**, 240 (1970); N. Jarmie, J.H. Jett, J.L. Detch, Jr., and R.L. Hutson, *ibid.* **25**, 34 (1970).
[40] J.D. Hutton, W. Haeberli, L.D. Knutson, and P. Signell, Phys. Rev. Lett. **35**, 429 (1975).
[41] P. Catillon, M. Chapellier, and D. Garreta, Nucl. Phys. **B2**, 93 (1967).
[42] S. Kikuchi, J. Sanada, S. Suwa, I. Hayashi, K. Nisimura, and K. Fukunaga, J. Phys. Soc. Jpn. **15**, 9 (1960).
[43] W.A. Blanpied, Phys. Rev. **116**, 738 (1959).
[44] K.W. Brockman, Jr., Phys. Rev. **110**, 163 (1958).
[45] N. Jarmie and J.H. Jett, Phys. Rev. C **13**, 2554 (1976).
[46] J.W. Burkig, J.R. Richardson, and G.E. Schrank, Phys. Rev. **113**, 290 (1959).
[47] H.N. Royden and B.T. Wright, Phys. Rev. **113**, 294 (1959).
[48] A. Abragam, M. Borghini, P. Catillon, J. Coustham, P. Roubeau, and J. Thirion, Phys. Lett. **2**, 310 (1962).
[49] P. Catillon, J. Sura, and A. Tarrats, Phys. Rev. Lett. **20**, 602 (1968).
[50] C.J. Batty, G.H. Stafford, and R.S. Gilmore, Nucl. Phys. **51**, 225 (1964).
[51] T.H. Jeong, L.H. Johnston, D.E. Young, and C.N. Waddell, Phys. Rev. **118**, 1080 (1960).
[52] N. Jarmie, J.E. Brolley, H. Kruse, H.C. Bryant, and R. Smythe, Phys. Rev. **155**, 1438 (1967).
[53] P. Christmas and A.E. Taylor, Nucl. Phys. **41**, 388 (1963).
[54] A. Ashmore, B.W. Davies, M. Devine, S.J. Hoey, J. Litt, M.E. Shepherd, R.C. Hanna, and L.P. Robertson, Nucl. Phys. **73**, 256 (1965).
[55] L.H. Johnston and Y.S. Tsai, Phys. Rev. **115**, 1293 (1959).
[56] C.J. Batty, R.S. Gilmore, and G.H. Stafford, Nucl. Phys. **45**, 481 (1963).
[57] D. Garreta, K. Nisimura, and M. Fruneau, Phys. Lett. **31B**, 363 (1970).
[58] L.H. Johnston and D.A. Swenson, Phys. Rev. **111**, 212 (1958).
[59] U.E. Kruse, J.M. Teem, and N.F. Ramsey, Phys. Rev. **101**, 1079 (1956).
[60] J.N. Palmieri, A.M. Cormack, N.F. Ramsey, and R. Wilson, Ann. Phys. (NY) **5**, 299 (1958).
[61] D. Garreta and M. Fruneau, Nucl. Phys. **A170**, 492 (1971).
[62] K. Nisimura, T. Hasegawa, T. Saito, E. Takasaki, N. Horikawa, T. Nakanishi, M. Saito, S. Sugimoto, T. Yamaki, and H. Ueno, Phys. Lett. **30B**, 612 (1969).

[63] A. Ashmore, M. Devine, B. Hird, J. Litt, W.H. Range, M.E. Shepherd, and R.L. Clarke, Nucl. Phys. **65**, 305 (1965).
[64] C.J. Batty, T.C. Griffith, D.C. Imrie, G.J. Lush, and L.A. Robbins, Nucl. Phys. **A98**, 489 (1967).
[65] T.C. Griffith, D.C. Imrie, G.J. Lush, and A.J. Metheringham, Phys. Rev. Lett. **10**, 444 (1963).
[66] A. Berdoz, F. Foroughi, and C. Nussbaum, J. Phys. G **12**, L133 (1986).
[67] K. Nisimura, J. Sanada, P. Catillon, K. Fukunaga, T. Hasegawa, H. Hasai, N. Ryu, D.C. Worth, and H. Imada, Prog. Theor. Phys. **30**, 719 (1963).
[68] J. Sanada, K. Kuriyama, Y. Takeuchi, Y. Nojiri, N. Ryu, H. Hasai, M. Ikeda, S. Kobayashi, K. Nagamine, D.C. Worth, and T. Yamaya, Nucl. Phys. **B4**, 379 (1968).
[69] N. Tamura, K. Imai, K. Nisimura, K. Egawa, K. Hatanaka, H. Shimizu, T. Saito, H. Sato, Y. Wakuta, H. Hasuyama, Y. Mori, and H. Yamamoto, J. Phys. Soc. Jpn. **44**, S289 (1978).
[70] D.E. Young and L.H. Johnston, Phys. Rev. **119**, 313 (1960).
[71] O.N. Jarvis, T.W.P. Brogden, B. Rose, J.P. Scanlon, J. Orchard-Webb, and M.R. Wigan, Nucl. Phys. **A108**, 63 (1968).
[72] M.R. Wigan, R.A. Bell, P.J. Martin, O.N. Jarvis, and J.P. Scanlon, Nucl. Phys. **A114**, 377 (1968).
[73] A.E. Taylor, E. Wood, and L. Bird, Nucl. Phys. **16**, 320 (1960).
[74] E.H. Thorndike and T.R. Ophel, Phys. Rev. **119**, 362 (1960).
[75] O.N. Jarvis, B. Rose, G.F. Cox, and G.H. Eaton, Nucl. Phys. **61**, 194 (1965).
[76] E. Baskir, E.M. Hafner, A. Roberts, and J.H. Tinlot, Phys. Rev. **106**, 564 (1957).
[77] S. Hee and R. Wilson, Phys. Rev. **132**, 2236 (1963).
[78] C. Caverzasio, A. Michalowicz, K. Kuroda, M. Poulet, and N. Poutcherov, J. Phys. Radium **24**, 1048 (1963).
[79] S. Hee and E.H. Thorndike, Phys. Rev. **132**, 744 (1963).
[80] O.N. Jarvis, B. Rose, G.H. Eaton, and C.P. van Zyl, Nucl. Phys. **50**, 529 (1964).
[81] E.H. Thorndike, J. Lefrançois, and R. Wilson, Phys. Rev. **120**, 1819 (1960).
[82] G.F. Cox, G.H. Eaton, C.P. van Zyl, O.N. Jarvis, and B. Rose, Nucl. Phys. **B4**, 353 (1967).
[83] L. Bird, D.N. Edwards, B. Rose, A.E. Taylor, and E. Wood, Nucl. Phys. **42**, 280 (1963).
[84] C.F. Hwang, T.R. Ophel, E.H. Thorndike, and R. Wilson, Phys. Rev. **119**, 352 (1960).
[85] L. Bird, P. Christmas, A.E. Taylor, and E. Wood, Nucl. Phys. **27**, 586 (1961).
[86] O.N. Jarvis, B. Rose, J.P. Scanlon, and E. Wood, Nucl. Phys. **42**, 294 (1963).
[87] O.N. Jarvis, C. Whitehead, and M. Shah, Phys. Lett. **36B**, 409 (1971).
[88] J.M. Cassels, Proc. Phys. Soc. London, Sect. A **69**, 495 (1956).
[89] A. Michalowicz, Rev. Mod. Phys. **39**, 523 (1967).
[90] O. Chamberlain and J.D. Garrison, Phys. Rev. **103**, 1860 (1956).
[91] D. Fischer and J. Baldwin, Phys. Rev. **100**, 1445 (1955).
[92] P. Hillman, A. Johansson, and G. Tibell, Phys. Rev. **110**, 1218 (1958).
[93] B.A. Ryan, A. Kanofsky, T.J. Devlin, R.E. Mischke, and P.F. Shepard, Phys. Rev. D **3**, 1 (1971).
[94] L.G. Greeniaus, D.A. Hutcheon, C.A. Miller, G.A. Moss, G. Roy, R. Dubois, C. Amsler, B.K.S. Koene, and B.T. Murdoch, Nucl. Phys. **A322**, 308 (1979).

[95] D. Axen, L. Felawka, S. Jaccard, J. Vávra, G.A. Ludgate, N.M. Stewart, C. Amsler, R.C. Brown, D.V. Bugg, J.A. Edgington, C.J. Oram, K. Shakarchi, and A.S. Clough, Lett. Nuovo Cimento **20**, 151 (1977).
[96] E.H. Thorndike, Rev. Mod. Phys. **39**, 513 (1967).
[97] A. Abashian and E.M. Hafner, Phys. Rev. Lett. **1**, 255 (1958).
[98] J.H. Tinlot and R.E. Warner, Phys. Rev. **124**, 890 (1961).
[99] J.F. Marshall, C.N. Brown, and F. Lobkowicz, Phys. Rev. **150**, 1119 (1966).
[100] K. Gotow, F. Lobkowicz, and E. Heer, Phys. Rev. **127**, 2206 (1962).
[101] A.C. England, W.A. Gibson, K. Gotow, E. Heer, and J. Tinlot, Phys. Rev. **124**, 561 (1961).
[102] K. Gotow and F. Lobkowicz, Phys. Rev. **136**, B1345 (1964).
[103] Y. Onel, R. Hausammann, E. Heer, R. Hess, C. Lechanoine-Leluc, W.R. Leo, D. Rapin, S. Jaccard, and S. Mango, Phys. Rev. D **40**, 35 (1989).
[104] C. Amsler, D. Axen, J. Beveridge, D.V. Bugg, A.S. Clough, R. Dubois, J.A. Edgington, D. Gibson, R. Keeler, G.A. Ludgate, C.J. Oram, J.R. Richardson, L.P. Robertson, and N.M. Stewart, J. Phys. G **4**, 1047 (1978).
[105] O. Chamberlain, E. Segrè, R.D. Tripp, C. Wiegand, and T. Ypsilantis, Phys. Rev. **105**, 288 (1957).
[106] D. Aebischer, B. Favier, L.G. Greeniaus, R. Hess, A. Junod, C. Lechanoine, J.C. Niklès, D. Rapin, C. Richard-Serre, and D.W. Werren, Phys. Rev. D **13**, 2478 (1976).
[107] D. Ottewell, P. Walden, E.G. Auld, G. Giles, G. Jones, G.J. Lolos, B.J. McParland, W. Ziegler, and W. Falk, Nucl. Phys. **A412**, 189 (1984).
[108] I.P. Auer, W.R. Ditzler, D. Hill, K. Imai, H. Spinka, R. Stanek, K. Toshioka, D. Underwood, R. Wagner, A. Yokosawa, E.W. Hoffman, J.J. Jarmer, G.R. Burleson, W.B. Cottingame, S.J. Greene, and S. Stuart, Phys. Rev. D **29**, 2435 (1984).
[109] A. Beretvas, Phys. Rev. **171**, 1392 (1968).
[110] D. Cheng, B. Macdonald, J.A. Helland, and P.M. Ogden, Phys. Rev. **163**, 1470 (1967).
[111] D. Besset, Q.H. Do, B. Favier, L.G. Greeniaus, R. Hess, C. Lechanoine, D. Rapin, D.W. Werren, and Ch. Weddigen, Phys. Rev. D **21**, 580 (1980).
[112] I.M. Vasilevskiĭ, V.V. Vishnyakov, É. Iliesku, and A.A. Tyapkin, Zh. Eksp. Teor. Fiz. **45**, 474 (1963) [Sov. Phys.–JETP **18**, 327 (1964)].
[113] Yu. M. Kazarinov, F. Lehar, G. Peter, A.F. Pisarev, and K.M. Vahlbruch, Zh. Eksp. Teor. Fiz. **47**, 848 (1964) [Sov. Phys.–JETP **20**, 565 (1965)].
[114] J.E. Simmons, Phys. Rev. **104**, 416 (1956).
[115] M.W. McNaughton and E.P. Chamberlin, Phys. Rev. C **24**, 1778 (1981).
[116] J.V. Allaby, A. Ashmore, A.N. Diddens, J. Eades, G.B. Huxtable, and K. Skarsvåg, Proc. Phys. Soc. London **77**, 234 (1961).
[117] F. Shimizu, H. Koiso, Y. Kubota, F. Sai, S. Sakamoto, and S.S. Yamamoto, Nucl. Phys. **A389**, 445 (1982).
[118] F. Betz, J. Arens, O. Chamberlain, H. Dost, P. Grannis, M. Hansroul, L. Holloway, C. Schultz, and G. Shapiro, Phys. Rev. **148**, 1289 (1966).
[119] K. Abe, B.A. Barnett, J.H. Goldman, A.T. Laasanen. P.H. Steinberg, G.J. Marmer, D.R. Moffett, and E.F. Parker, Phys. Rev. D **12**, 1 (1975).
[120] M.G. Albrow, S. Andersson/Almehed, B. Bošnjaković, C. Daum, F.C. Erné, J.P. Lagnaux, J.C. Sens, and F. Udo, Nucl. Phys. **B23**, 445 (1970).

[121] J. Bystricky and F. Lehar, *Nucleon-Nucleon Scattering Data*, Vol. 11-1 of *Physics Data*, edited by H. Behrens and G. Ebel (Fachinformationszentrum, Karlsruhe, Federal Republic of Germany, 1978).
[122] J. Bystricky and F. Lehar, *Nucleon-Nucleon Scattering Data*, Vol. 11-3 of *Physics Data*, edited by H. Behrens and G. Ebel (Fachinformationszentrum, Karlsruhe, Federal Republic of Germany, 1981).
[123] Scattering Analysis Interactive Dial-up (SAID), Virginia Polytechnic Institute, Blacksburg, Virginia (R.A. Arndt, private communication).
[124] M.S. Sher, P. Signell, and L. Heller, Ann. Phys. (NY) **58**, 1 (1970).
[125] W.A. van der Sanden, A.H. Emmen, and J.J. de Swart, Institution Report No. THEF-NIJM 83.11, Nijmegen, 1983 (unpublished).
[126] O.N. Jarvis and B. Rose, Phys. Lett. **15**, 271 (1965).
[127] J. Bystricky, F. Lehar, and P. Winternitz, J. Phys. (Paris) **39**, 1 (1978).
[128] M.H. MacGregor, R.A. Arndt, and R.M. Wright, Phys. Rev. **169**, 1128 (1968).
[129] M.H. MacGregor, R.A. Arndt, and R.M. Wright, Phys. Rev. **182**, 1714 (1969).
[130] R.A. Arndt, R.H. Hackman, and L.D. Roper, Phys. Rev. C **9**, 555 (1974).
[131] R.A. Arndt, R.H. Hackman, and L.D. Roper, Phys. Rev. C **15**, 1002 (1977).
[132] J.L. Yntema and M.G. White, Phys. Rev. **95**, 1226 (1954).
[133] P. Kroll, *Phenomenological Analysis of Nucleon-Nucleon Scattering*, Vol. 22-1 of *Physics Data*, edited by H. Behrens and G. Ebel (Fachinformationszentrum, Karlsruhe, Federal Republic of Germany, 1981).
[134] D.V. Bugg, Nucl. Phys. **B5**, 29 (1968).
[135] G. Breit, M. Tischler, S. Mukherjee, and J. Lucas, Proc. Nat. Acad. Sci. USA **68**, 897 (1971).
[136] V.G.J. Stoks, P.C. van Campen, W. Spit, and J.J. de Swart, Phys. Rev. Lett. **60**, 1932 (1988).
[137] V.G.J. Stoks, P.C. van Campen, T.A. Rijken, and J.J. de Swart, Phys. Rev. Lett. **61**, 1702 (1988).

## TABLES

TABLE I. Data reference table. A dagger denotes data not included in the Nucleon-Nucleon Scattering Data Tables (Refs. [121, 122]). An asterisk denotes data not included in the data set NN896 of SAID (Ref. [123]).

| $T_{\rm lab}$ (MeV) | No.[a], type | $\chi^2$ | %error | Pred. norm[b] | Reject | Reference | Comment | Institute |
|---|---|---|---|---|---|---|---|---|
| 0.33766–0.40517 | 6 $\sigma$ | 3.852 | float[c] | 1.3222 | 0.372 83 MeV | [27] | d | Los Alamos |
| 0.35003–0.42003† | 36 $\sigma$ | 39.194 | float[c] | 0.9975 | | [28] | e | Zürich |
| 0.35009† | 17 $\sigma$ | 25.172 | 0.16 | 0.9993 | | [28] | e | Zürich |
| 0.40004† | 3 $\sigma$ | 1.033 | 0.21 | 1.0009 | | [28] | | Zürich |
| 0.42006† | 22 $\sigma$ | 38.545 | 0.16 | 0.9995 | | [28] | e | Zürich |
| 0.49923† | 39 $\sigma$ | 32.020 | 0.16 | 0.9989 | | [28] | e | Zürich |
| 0.49925†* | 3 $\sigma$ | | 0.09 | | all | [29] | f,g | Basel |
| 0.74996† | 26 $\sigma$ | 16.825 | 0.16 | 0.9987 | | [28] | e | Zürich |
| 0.99183† | 31 $\sigma$ | 25.475 | 0.16 | 0.9989 | | [28] | e | Zürich |
| 0.9919†* | 3 $\sigma$ | | 0.09 | | all | [29] | f,g | Basel |
| 0.9919†* | 2 $\sigma$ | | float[c] | | all | [30] | f,g | Basel |
| 1.397 | 11 $\sigma$ | | none | | all | [31] | f | Wisconsin |
| 1.855 | 13 $\sigma$ | | none | | all | [31] | f | Wisconsin |
| 1.8806†* | 3 $\sigma$ | | 0.09 | | all | [29] | f,g | Basel |
| 2.425 | 14 $\sigma$ | | none | | all | [31] | f | Wisconsin |
| 3.037 | 13 $\sigma$ | | none | | all | [31] | f | Wisconsin |
| 4.978 | 17 $\sigma$ | 17.272 | 0.4 | 1.0029 | | [32] | | Kyoto |
| 5.05† | 11 $P$ | 5.284 | 1.0 | 0.9980 | | [33] | | Wisconsin |
| 6.141†* | 17 $\sigma$ | | 0.4 | | all | [34] | h,i | Berkeley |
| 6.141†* | 6 $P$ | | none | | all | [35] | j | Erlangen |
| 6.968 | 17 $\sigma$ | 15.922 | 0.4 | 1.0051 | | [32] | | Kyoto |
| 8.030 | 17 $\sigma$ | 12.128 | 0.4 | 1.0051 | | [32] | | Kyoto |
| 8.097 | 16 $\sigma$ | | 0.4 | | all | [34] | h,i | Berkeley |
| 9.57† | 1 $A_{yy}$ | 0.234 | none | | | [36] | | Erlangen |
| 9.6* | 5 $P$ | 2.172 | none | | | [37] | k | Berkeley |
| 9.69 | 26 $\sigma$ | 18.651 | 0.73 | 0.9853 | | [38] | l,m | Minnesota |
| 9.69 | 5 $\sigma$ | 5.476 | (0.36) | 0.9819 | | [39] | n | Los Alamos |
| 9.85† | 15 $P$ | 17.381 | 1.0 | 0.9956 | | [33] | | Wisconsin |
| 9.918* | 17 $\sigma$ | | 0.4 | | all | [34] | h,i | Berkeley |
| 9.918 | 11 $\sigma$ | 19.902 | 0.38 | 0.9959 | 20.05° | [39] | | Los Alamos |
| 10.00 | 7 $P$ | 9.684 | none | | | [40] | | Wisconsin |
| 11.40 | 1 $A_{yy}/A_{xx}$ | 0.083 | none | | | [41] | | Saclay |
| 13.60 | 11 $\sigma$ | 15.468 | 0.33 | 1.0024 | | [39] | | Los Alamos |
| 14.16 | 17 $\sigma$ | | none | | all | [42] | j,l | Tokyo |
| 15.6* | 5 $P$ | 1.679 | none | | 36.14° | [37] | k | Berkeley |
| 16.2 | 1 $P$ | 0.679 | none | | | [43] | l | Princeton |
| 17.7* | 1 $P$ | 0.328 | none | | | [44] | | Princeton |
| 19.15–26.50 | 3 $A_{xx}$, 3 $A_{yy}$ | 2.484 | float[c] | 1.013(8) | | [41] | | Saclay |
| 19.7†* | 13 $\sigma$ | 8.482 | 0.37 | 0.9987 | | [45] | | Los Alamos |
| 19.7* | 6 $P$ | | none | | all | [37] | k,o | Berkeley |
| 19.8* | 15 $\sigma$ | 22.039 | none | | | [46] | | UCLA |
| 19.84* | 7 $\sigma$ | 8.608 | none | | | [47] | p | UCLA |
| 20.00* | 1 $C_{nn}$ | 1.131 | none | | | [48] | | Saclay |
| 20.20 | 8 $P$ | | 12.0 | | all | [49] | j | Saclay |
| 21.95–50.02 | 7 $\sigma$ | | 0.36 | | all | [50] | o | Rutherford |
| 25.63 | 23 $\sigma$ | | 0.93 | | all | [51] | j | Minnesota |
| 27.05 | 1 $C_{nn}$ | 0.309 | none | | | [52] | | Los Alamos |
| 27.4* | 1 $P$ | 0.144 | none | | | [53] | q | Harwell |
| 27.6 | 3 $R$, 3 $A$ | 13.078 | 3.0 | 1.035(29) | | [54] | | Rutherford |
| 28.16 | 1 $\sigma$ | 0.731 | none | | | [55] | l | Minnesota |
| 30.0 | 2 $P$ | 4.498 | 4.0 | 1.0340 | | [56] | | Rutherford |
| 31.15 | 1 $\sigma$ | 0.002 | none | | | [55] | l | Minnesota |
| 34.20 | 1 $\sigma$ | 0.886 | none | | | [55] | l | Minnesota |
| 36.8* | 2 $P$ | 0.417 | none | | | [53] | q | Harwell |
| 36.90 | 1 $\sigma$ | 0.198 | none | | | [55] | l | Minnesota |
| 37.23 | 1 $A_{xx}$, 1 $A_{yy}$ | 1.123 | float[c] | 1.0325 | | [57] | | Saclay |
| 38.3* | 1 $P$ | 2.997 | none | | | [53] | q | Harwell |
| 39.4 | 27 $\sigma$ | 31.327 | 0.93 | 1.0080 | | [58] | l | Minnesota |
| 39.60 | 1 $\sigma$ | 0.209 | none | | | [55] | l | Minnesota |

| | | | | | | | | |
|---|---|---|---|---|---|---|---|---|
| 41.0 | 1 $\sigma$ | 0.814 | none | | | [59] | | Harvard |
| 44.66 | 1 $\sigma$ | 1.249 | none | | | [55] | l | Minnesota |
| 46.0–86.0* | 5 $\sigma$ | | float[c] | | all | [60] | o | Harvard |
| 46.0 | 1 $P$ | 1.244 | 2.8 | 1.0017 | | [60] | r | Harvard |
| 46.9†* | 1 $A_{xx}$, 1 $A_{yy}$ | 0.025 | float[c] | 1.064(17) | | [61] | | Saclay |
| 47.5 | 1 $A_{xx}$, 1 $A_{yy}$ | 2.673 | float[c] | 1.2531 | | [62] | | Kyoto |
| 47.5 | 5 $A$ | 7.123 | 5.0 | 1.0178 | | [63] | | Rutherford |
| 47.8 | 5 $R$ | 5.116 | 5.0 | 0.9977 | | [54] | | Rutherford |
| 47.8 | 5 $A$ | 2.434 | 5.0 | 0.9809 | | [54] | | Rutherford |
| 49.41 | 30 $\sigma$ | 31.164 | 0.32 | 1.0011 | 14.02°, 17.06° | [64] | | Rutherford |
| 49.7* | 1 $P$ | 8.257 | none | | | [53] | q | Harwell |
| 49.9 | 2 $P$ | 0.008 | 3.0 | 1.0005 | | [56] | | Rutherford |
| 50.0 | 1 $D$ | 1.534 | none | | | [65] | | Harwell |
| 50.06†* | 24 $\sigma$ | 12.823 | 1.6 | 0.97820 | | [66] | | SIN |
| 50.17 | 1 $\sigma$ | 0.487 | none | | | [55] | l | Minnesota |
| 51.5 | 1 $\sigma$ | 0.869 | none | | | [59] | | Harvard |
| 51.7 | 1 $P$ | 0.482 | none | | | [53] | q | Harwell |
| 52.0 | 1 $A_{yy}$, 1 $C_{kp}$ | 5.406 | none | | | [67] | | Tokyo |
| 52.34 | 29 $\sigma$ | 23.663 | 0.53 | 1.0037 | 16.22° | | | |
| | | | | | 18.24°, 20.27° | [68] | | Tokyo |
| 52.34† | 12 $P$ | 7.643 | 2.0 | 0.9846 | | [69] | | Kyoto |
| 53.2 | 1 $P$ | 1.412 | none | | | [53] | q | Harwell |
| 56.0 | 1 $P$ | 0.317 | 2.8 | 1.0032 | | [60] | r | Harvard |
| 56.15 | 1 $\sigma$ | 0.241 | none | | | [55] | l | Minnesota |
| 58.5 | 1 $P$ | 0.538 | none | | | [53] | q | Harwell |
| 61.92 | 1 $\sigma$ | 0.012 | none | | | [55] | l | Minnesota |
| 66.0 | 11 $\sigma$ | | float[c] | | all | [60] | j | Harvard |
| 66.0 | 11 $P$ | 8.323 | 2.8 | 0.9769 | | [60] | r | Harvard |
| 68.19† | 12 $P$ | 6.533 | 2.0 | 0.9826 | | [69] | | Kyoto |
| 68.3 | 26 $\sigma$ | 30.733 | 1.1 | 0.9923 | | [70] | l | Minnesota |
| 68.42 | 1 $\sigma$ | 0.006 | none | | | [55] | l | Minnesota |
| 69.5 | 1 $\sigma$ | 0.037 | none | | | [59] | | Harvard |
| 70.0* | 5 $\sigma$ | 9.471 | float[c] | 5.7558 | 25.0° | [59] | | Harvard |
| 70.0 | 1 $P$ | 2.031 | none | | | [53] | q | Harwell |
| 71.0 | 1 $P$ | 0.465 | 2.8 | 1.0040 | | [60] | r | Harvard |
| 73.5 | 1 $A_{yy}$ | 0.704 | none | | | [71] | | Harwell |
| 78.0 | 1 $P$ | 0.552 | 2.8 | 0.9935 | | [60] | r | Harvard |
| 78.5 | 1 $\sigma$ | 0.003 | none | | | [59] | | Harvard |
| 86.0 | 1 $P$ | 0.014 | 2.8 | 1.0012 | | [60] | r | Harvard |
| 95.0 | 1 $\sigma$ | 0.016 | none | | | [59] | | Harvard |
| 95.0 | 6 $\sigma$ | 2.318 | none | | | [59] | | Harvard |
| 95.0 | 5 $\sigma$ | 3.820 | float[c] | 4.6381 | 25.0° | [59] | | Harvard |
| 95.0 | 14 $\sigma$ | 5.742 | float[c] | 1.0287 | | [60] | | Harvard |
| 95.0 | 14 $P$ | 13.445 | 2.8 | 1.0082 | | [60] | r | Harvard |
| 97.0 | 1 $P$ | 2.329 | none | | | [53] | q | Harwell |
| 97.7 | 13 $P$ | 14.303 | 0.85 | 1.0092 | | [72] | | Harwell |
| 98.0* | 14 $\sigma$ | | 4.5 | | all | [73] | o | Harwell |
| 98.0 | 14 $P$ | 13.287 | 2.2 | 1.0225 | | [73] | s | Harwell |
| 98.0* | 5 $P$ | 4.634 | 3.0 | 1.0058 | | [74] | t | Harvard |
| 98.0* | 3 $P$ | 6.273 | 3.0 | 1.0365 | | [74] | u | Harvard |
| 98.0 | 5 $D$ | 7.596 | 3.0 | 1.0046 | | [74] | | Harvard |
| 98.0 | 1 $A_{yy}$ | 0.013 | none | | | [71] | | Harwell |
| 98.1 | 5 $R$ | 6.424 | none | | | [75] | | Harwell |
| 98.1 | 4 $R, R'$ | 1.168 | none | | | [75] | v | Harwell |
| 98.8 | 19 $\sigma$ | 16.399 | 1.0 | 0.9948 | | [72] | | Harwell |
| 102.0 | 3 $\sigma$ | 7.565 | float[c] | 1.0122 | | [60] | | Harvard |
| 102.0 | 3 $P$ | 1.940 | 2.8 | 0.9816 | | [60] | r | Harvard |
| 107.0 | 3 $\sigma$ | 1.167 | float[c] | 1.0722 | | [60] | | Harvard |
| 107.0* | 3 $P$ | 7.234 | 2.8 | 1.0146 | | [60] | r | Harvard |
| 118.0 | 16 $\sigma$ | 23.673 | float[c] | 1.0448 | 86.8° | [60] | | Harvard |
| 118.0* | 16 $P$ | 31.718 | 2.8 | 1.0507 | | [60] | r | Harvard |
| 127.0 | 3 $\sigma$ | 0.766 | float[c] | 1.0400 | | [60] | | Harvard |
| 127.0* | 3 $P$ | 8.247 | 2.8 | 1.0171 | | [60] | r | Harvard |
| 130.0 | 4 $P$ | 3.059 | 3.3 | 1.0197 | | [76] | | Rochester |
| 137.0 | 3 $\sigma$ | 0.023 | float[c] | 1.0120 | | [60] | | Harvard |
| 137.0 | 3 $P$ | 3.524 | 2.8 | 1.0279 | | [60] | r | Harvard |
| 137.5 | 5 $R'$ | 1.123 | 5.0 | 0.9993 | | [77] | | Harvard |
| 138.0 | 17 $P$ | 25.717 | none | | 82°, 86° | [78] | w | Orsay |

| | | | | | | | | |
|---|---|---|---|---|---|---|---|---|
| 138.0* | 4 $D$ | 3.986 | none | | | [78] | w | Orsay |
| 139.0 | 6 $A$ | 3.929 | 4.0 | 1.0439 | | [79] | | Harvard |
| 140.4 | 6 $R, R'$ | 6.755 | none | | 72.1° | [80] | v | Harwell |
| 140.5 | 6 $R$ | 4.263 | none | | | [81] | x | Harvard |
| 140.7 | 20 $P$ | 19.053 | 0.85 | 0.9970 | | [82] | | Harwell |
| 141.0 | 8 $R$ | 10.139 | none | | | [83] | | Harwell |
| 142.0* | 20 $\sigma$ | | 4.5 | | all | [73] | o | Harwell |
| 142.0 | 29 $P$ | 29.356 | 2.2 | 0.9544 | 6.23°, 78.05° | [73] | s | Harwell |
| 142.0* | 8 $P$ | | 3.0 | | all | [84] | u,o | Harvard |
| 142.0* | 8 $P$ | 7.049 | 3.0 | 1.0016 | | [84] | t | Harvard |
| 142.0 | 8 $D$ | 11.279 | 3.0 | 1.0008 | | [84] | | Harvard |
| 143.0 | 7 $P$ | 5.102 | none | | | [85] | | Harwell |
| 143.0 | 7 $D$ | 7.444 | none | | | [85] | | Harwell |
| 143.0 | 6 $A$ | 4.886 | none | | | [86] | | Harwell |
| 143.2 | 1 $P$ | 1.881 | none | | | [71] | | Harwell |
| 143.2 | 2 $A_{yy}$ | 0.304 | none | | | [71] | | Harwell |
| 144.0 | 27 $\sigma$ | 36.142 | 0.58 | 0.9936 | 7.26° | [87] | | Harwell |
| 144.1 | 6 $\sigma$ | 3.042 | 0.88 | 0.9944 | | [82] | | Harwell |
| 144.1 | 15 $\sigma$ | 31.931 | 0.56 | 0.9861 | | [82] | | Harwell |
| 147.0* | 3 $\sigma$ | 1.332 | float[c] | 0.9955 | | [60] | | Harvard |
| 147.0* | 12 $\sigma$ | 20.947 | float[c] | 0.8526 | 4.13° | [60] | | Harvard |
| 147.0* | 15 $\sigma$ | 9.151 | float[c] | 0.9263 | | [60] | | Harvard |
| 147.0* | 11 $\sigma$ | 7.360 | float[c] | 0.9251 | | [60] | | Harvard |
| 147.0* | 1 $\sigma$ | 1.004 | none | | | [88] | | Harwell |
| 147.0 | 3 $P$ | 1.635 | 2.8 | 0.9946 | | [60] | r | Harvard |
| 147.0 | 12 $P$ | 16.225 | 2.8 | 1.0042 | | [60] | r | Harvard |
| 147.0 | 15 $P$ | 14.942 | 2.8 | 0.9933 | | [60] | r | Harvard |
| 147.0 | 13 $P$ | 22.975 | 2.8 | 0.9908 | | [60] | r | Harvard |
| 156.0 | 24 $\sigma$ | 24.692 | none | | 8.3°, 25.0° | [89] | | Orsay |
| 170.0 | 6 $\sigma$ | 10.412 | 6.9 | 1.0674 | 10.1° | [90] | | Berkeley |
| 170.0 | 5 $P$ | 3.450 | 6.6 | 1.0510 | | [91] | | Berkeley |
| 170.0 | 7 $P$ | 2.760 | 3.3 | 1.0154 | | [76] | | Rochester |
| 174.0* | 7 $\sigma$ | | 6.9 | | all | [90] | o | Berkeley |
| 176.0–179.0* | 2 $P$ | 1.452 | 3.7 | 0.9769 | | [92] | | Uppsala |
| 203.6* | 3 $\sigma$ | 10.778 | 5.4 | 1.0138 | | [93] | | Princeton |
| 205.0 | 1 $P$ | 1.705 | none | | | [94] | | TRIUMF |
| 209.1 | 3 $P$ | 2.578 | 2.8 | 0.9924 | | [95] | y | TRIUMF |
| 209.1 | 3 $D$ | 8.160 | 2.8 | 0.9782 | | [95] | y | TRIUMF |
| 209.1 | 3 $R$ | 0.559 | 2.8 | 0.9962 | | [95] | y | TRIUMF |
| 209.1 | 1 $R'$ | 1.700 | 2.8 | 1.0181 | | [95] | y | TRIUMF |
| 210.0† | 7 $\sigma$ | 6.855 | 1.64 | 1.0174 | | [96] | | Rochester |
| 210.0* | 2 $P$ | 2.680 | none | | | [97] | z | Rochester |
| 210.0* | 9 $P$ | 8.747 | 3.3 | 0.9912 | 67.8° | [76] | aa | Rochester |
| 210.0 | 7 $P$ | 1.905 | 3.6 | 1.0065 | | [98] | bb | Rochester |
| 213.0 | 13 $\sigma$ | 10.576 | 1.3 | 0.9978 | 8.9°, 9.8° | [99] | | Rochester |
| 213.0 | 13 $P$ | 23.508 | 3.1 | 1.0030 | | [99] | cc | Rochester |
| 213.0* | 7 $P$ | 14.249 | none | | 90.0° | [100] | dd | Rochester |
| 213.0* | 7 $P$ | 12.001 | none | | 90.0° | [100] | dd | Rochester |
| 213.0* | 7 $P$ | 8.664 | none | | 90.0° | [100] | dd | Rochester |
| 213.0 | 7 $D$ | 6.637 | none | | | [100] | dd | Rochester |
| 213.0 | 7 $R$ | 2.846 | none | | | [101] | ee | Rochester |
| 213.0 | 5 $R, A$ | 3.111 | none | | | [101] | ff | Rochester |
| 213.0 | 2 $A$ | 1.331 | none | | | [96] | | Rochester |
| 213.0 | 7 $R, R'$ | 13.364 | none | | 60.0° | [102] | v | Rochester |
| 217.0* | 7 $P$ | 12.719 | 2.2 | 1.0360 | | [98] | | Rochester |
| 225.0 | 1 $P$ | 0.932 | none | | | [94] | | Rochester |
| 241.0† | 8 $P$ | 18.797 | 3.0 | 0.9878 | | [103] | | SIN |
| 241.0† | 8 $D$ | 18.405 | 4.0 | 0.9470 | | [103] | | SIN |
| 241.0† | 8 $D_t$ | | 7.0 | | all | [103] | j | SIN |
| 241.0† | 8 $R, R'$ | 3.319 | 4.0 | 0.9586 | | [103] | v | SIN |
| 241.0† | 8 $A, A'$ | 13.003 | 4.0 | 0.9982 | | [103] | gg | SIN |
| 241.0† | 8 $M_{\omega 0sn}$ | 9.274 | 8.0 | 1.0144 | | [103] | hh | SIN |
| 241.0† | 8 $M_{\omega 0kn}$ | 8.511 | 8.0 | 0.9847 | | [103] | ii | SIN |
| 259.0 | 6 $\sigma$ | | 5.2 | | all | [90] | o | Berkeley |
| 260.0* | 6 $\sigma$ | 8.511 | 5.2 | 0.9867 | 9.3° | [90] | | Berkeley |
| 266.0 | 1 $P$ | 0.483 | 0.4 | 1.0006 | | [104] | | TRIUMF |
| 276.0 | 6 $P$ | 8.064 | 7.5 | 1.1213 | | [105] | | Berkeley |

| | | | | | | | | |
|---|---|---|---|---|---|---|---|---|
| 285.0 | 22 $\sigma$ | 25.886 | float[c] | 1.0348 | | [106] | | CERN |
| 300.0† | 1 $\sigma$ | 2.448 | 1.8 | 0.9737 | | [107] | | TRIUMF |
| 302.9† | 10 $A_{zz}$ | 15.307 | 2.0 | 0.9888 | | [108] | jj | Argonne |
| 305.0 | 14 $C_{nn}$ | 16.864 | 8.0 | 1.1432 | | [109] | | Chicago |
| 307.0 | 7 $P$ | 8.491 | 3.0 | 1.0050 | 83.7° | [110] | | Berkeley |
| 308.0 | 1 $P$ | 0.013 | none | | | [94] | | TRIUMF |
| 310.0 | 6 $D$ | 6.881 | none | | | [105] | | Berkeley |
| 310.0 | 6 $R$ | 9.778 | none | | | [105] | | Berkeley |
| 312.0 | 13 $P$ | 19.372 | 0.6 | 1.0018 | | [111] | | CERN |
| 312.0 | 13 $D$ | 7.849 | 1.2 | 0.9987 | | [111] | | CERN |
| 312.0 | 13 $R$ | 13.706 | 1.2 | 0.9997 | | [111] | | CERN |
| 312.0 | 13 $A$ | 7.729 | 1.2 | 1.0012 | 4.32° | [111] | | CERN |
| 314.0† | 8 $P$ | 17.414 | 3.0 | 0.9381 | | [103] | | SIN |
| 314.0† | 8 $D_t$ | 2.002 | 7.0 | 1.0357 | | [103] | | SIN |
| 314.0† | 8 $R, R'$ | 7.786 | 4.0 | 0.9520 | | [103] | v | SIN |
| 314.0† | 8 $A, A'$ | 10.176 | 4.0 | 1.0797 | | [103] | gg | SIN |
| 314.0† | 8 $M_{\omega 0sn}$ | 3.992 | 8.0 | 0.9977 | | [103] | hh | SIN |
| 314.0† | 8 $M_{\omega 0kn}$ | 2.611 | 8.0 | 0.9842 | | [103] | ii | SIN |
| 315.0 | 7 $\sigma$ | 6.086 | 10.0 | 1.0282 | | [105] | | Berkeley |
| 315.0 | 7 $P$ | 8.142 | 4.0 | 1.0743 | | [105] | | Berkeley |
| 315.0 | 1 $A_{yy}$ | 0.414 | none | | | [112] | | Dubna |
| 315.0 | 1 $A_{yy}$, 1 $C_{kp}$ | 0.039 | none | | | [113] | | Dubna |
| 316.0 | 3 $A$ | 0.580 | none | | | [114] | | Berkeley |
| 318.0† | 1 $P$ | 0.075 | 1.0 | 0.9982 | | [115] | | Los Alamos |
| 320.0 | 1 $P$ | 0.002 | 0.4 | 1.0000 | | [104] | | TRIUMF |
| 320.0 | 1 $C_{nn}$ | 1.158 | none | | | [116] | | Liverpool |
| 324.0–327.8 | 4 $P$ | 1.742 | 1.5 | 0.9924 | | [95] | y | TRIUMF |
| 324.0 | 3 $P$ | 2.341 | 1.5 | 0.9997 | | [95] | y | TRIUMF |
| 324.0 | 3 $D$ | 3.185 | 1.5 | 0.9904 | | [95] | y | TRIUMF |
| 324.0 | 3 $R$ | 1.332 | 1.5 | 1.0042 | | [95] | y | TRIUMF |
| 324.0 | 1 $R'$ | 3.715 | 1.5 | 1.0084 | | [95] | y | TRIUMF |
| 325.0† | 19 $\sigma$ | 25.087 | 10.0 | 1.0201 | | [117] | | Tokyo |
| 327.0 | 1 $P$ | 0.676 | none | | | [94] | | TRIUMF |
| 328.0 | 5 $\sigma$ | 8.784 | 2.9 | 0.9635 | | [93] | | Princeton |
| 328.0 | 14 $P$ | 6.667 | 6.2 | 1.0201 | 85.3° | [118] | | Berkeley |
| 330.0 | 13 $C_{nn}$ | 4.675 | 8.0 | 1.0705 | | [109] | | Chicago |
| 332.5 | 12 $\sigma$ | 17.339 | (3.0) | 1.1358 | | [119] | n | Argonne |
| 334.5 | 10 $\sigma$ | 7.381 | none | | | [120] | | CERN |
| 334.5 | 11 $P$ | 11.323 | 6.0 | 1.0642 | | [120] | | CERN |
| 341.0† | 16 $P$ | 20.506 | 3.0 | 0.9978 | | [103] | | SIN |
| 341.0† | 16 $D$ | 20.244 | 4.0 | 0.9521 | | [103] | | SIN |
| 341.0† | 16 $R$ | 11.661 | 4.0 | 0.9659 | 74.0° | [103] | | SIN |
| 341.0† | 16 $A$ | 20.871 | 4.0 | 0.9604 | | [103] | | SIN |
| 348.0 | 22 $\sigma$ | 20.343 | float[c] | 1.0092 | | [106] | | CERN |
| 350.0† | 1 $\sigma$ | 1.027 | 1.8 | 1.0179 | | [107] | | TRIUMF |

[a]The number includes all published data.

[b]Predicted norm with which the experimental values should be multiplied before comparison with the theoretical values.

[c]Floated normalization because these data are relative only.

[d]0.37283 MeV rejected. This is in accordance with the result of the phase-shift analysis of Sher, Signell and Heller [124].

[e]Two extra angle-dependent normalizations included (see Ref. [28]).

[f]Rejected as a result of the analysis of van der Sanden, Emmen, and de Swart [125] of the 0–3 MeV data.

[g]The authors of Ref. [28] call these data doubtful, mainly because of the use of the current integration method, which they found unsatisfactory at low energies.

[h]Disagreement between single-group fit and multienergy fit too large.

[i]We use the BGS data (see Ref. [34]).

[j]Group rejected due to improbably low $\chi^2$ (rejection criteria).

[k]Numerical values were taken from the Nucleon-Nucleon Scattering Data Tables [121].
[l]Probable errors changed to standard errors ($\sigma \approx 1.48$ probable errors).
[m]Normalization error extracted from absolute error and relative error (see Ref. [38]).
[n]Floated normalization because norm contributes more than 9 to $\chi^2$.
[o]Group rejected due to improbably high $\chi^2$ (rejection criteria).
[p]Overall error taken (see Ref. [47]).
[q]Renormalized according to Jarvis and Rose [126]): polarization times 0.89, error unchanged.
[r]Renormalized according to Jarvis and Rose [126]): polarization times 0.933, error unchanged.
[s]Renormalized according to Jarvis and Rose [126]): polarization times 0.911, error unchanged.
[t]Analyzing power measurement.
[u]Asymmetry measurement.
[v]Measured was $R\cos(\theta) + R'\sin(\theta)$.
[w]Renormalized by a factor of 0.85 according to Michalowicz [89].
[x]Total error taken (see Ref. [81]).
[y]Numerical values from Bugg and co-workers [19].
[z]Measurement of polarization and analyzing power as a check for time-reversal invariance. We do not discriminate between them.
[aa]We use the C–$H_2$ data in accordance with the data at 130 and 170 MeV.
[bb]Renormalized according to Thorndike [96]: polarization times (0.985±0.035).
[cc]Renormalized according to Thorndike [96]: polarization times (0.970±0.030).
[dd]Datum at 90° not included, as recommended by the authors of this measurement [100]).
[ee]Revised values from Thorndike [96]).
[ff]Measured was $A\cos(\theta) + R\sin(\theta)$.
[gg]Measured was $A\cos(\theta) + A'\sin(\theta)$.
[hh]Measured was $\cos(\omega)M_{s'0sn} - \sin(\omega)M_{k'0sn}$ in the four-index notation as used in Ref. [127].
[ii]Measured was $\cos(\omega)M_{s'0kn} - \sin(\omega)M_{k'0kn}$ in the four-index notation as used in Ref. [127].
[jj]Normalization error from error in beam polarization and target polarization [108].

TABLE II. The neutral pion-nucleon coupling constant $10^3 \times f_0^2$.

| | | |
|---|---|---|
| MacGregor *et al.* (Ref. [128]) | 1968 | 81.4±4.6 |
| Bugg (Ref. [134]) | 1968 | 75.2±3.9 |
| Breit *et al.* (Ref. [135]) | 1971 | 73.1–81.8 |
| Bugg *et al.* (Ref. [19]) | 1978 | 77.8±3.6 |
| Kroll (Ref. [133]) | 1981 | 80.3±2.2 |
| Present result | 1989 | 74.9±0.7 |

TABLE III. $P$-matrix parameters for the parametrized waves (see text).

| | $c_0$ | $c_1$ | $c_2$ | $c_3$ | $c_4$ | $c_5$ | $c_6$ |
|---|---|---|---|---|---|---|---|
| $^1S_0$ | –0.853 | 0.185 | –0.201 | 0.390 | –0.194 | 0.040 | –0.003 |
| $^3P_0$ | 2.782 | –0.322 | 0.103 | 0 | 0 | 0 | 0 |
| $^3P_1$ | 2.738 | 0.783 | –0.248 | 0.033 | 0 | 0 | 0 |
| $^1D_2$ | 1.348 | –0.363 | –0.014 | 0 | 0 | 0 | 0 |
| $P_1(J=2)$ | –1.123 | –0.024 | –0.014 | 0 | 0 | 0 | 0 |
| $\theta(J=2)$ | 0.011 | 0.014 | 0 | 0 | 0 | 0 | 0 |
| $P_2(J=2)$ | 1.334 | 0 | 0 | 0 | 0 | 0 | 0 |
| $^3F_3$ | –0.825 | 0 | 0 | 0 | 0 | 0 | 0 |
| $P_1(J=4)$ | –4.700 | 0.885 | 0 | 0 | 0 | 0 | 0 |
| $\theta(J=4)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $P_2(J=4)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 |
| $^1G_4$ | 15.097 | 0 | 0 | 0 | 0 | 0 | 0 |

TABLE IV. SE results at ten energy clusters. $N_{\rm obs}$: number of scattering observables. $N_{\rm df}$: number of degrees of freedom. The phase parameters in degrees are with respect to EM wave functions and the SE error is the square root of the diagonal elements of the full error matrix. $\delta^C_{C+EM}$ denotes the EM phase shift.

| 0.38254 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
|---|---|---|---|---|
| | 122 | 118 | 139.82 | 134.09 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 14.6075 | 14.6076 | 0.0025 | –0.0985 |
| $\Delta_C$ | –0.0009 | –0.0069 | 0.0025 | –0.0547 |
| | | | | |
| 1.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
| | 57 | 55 | 42.30 | 38.08 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 32.6667 | 32.6813 | 0.0094 | –0.0805 |
| $\Delta_C$ | –0.0045 | –0.0096 | 0.0029 | –0.0503 |
| | | | | |
| 5.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
| | 45 | 40 | 38.48 | 30.74 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 54.7423 | 54.5464 | 0.0871 | –0.0376 |
| $^3P_0$ | 1.5793 | 1.6144 | 0.0927 | –0.1248 |
| $^3P_1$ | –0.8961 | –0.9072 | 0.0282 | –0.0427 |
| $^3P_2$ | 0.2129 | 0.2272 | 0.0177 | –0.0153 |
| $^1D_2$ | 0.0433 | 0.0395 | 0.0096 | –0.0282 |
| | | | | |
| 10.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
| | 103 | 97 | 103.54 | 86.68 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 55.1273 | 55.0896 | 0.0681 | –0.0156 |
| $^3P_0$ | 3.7222 | 3.5426 | 0.0729 | –0.1546 |
| $^3P_1$ | –2.0442 | –2.1048 | 0.0266 | –0.0387 |
| $^3P_2$ | 0.6462 | 0.6601 | 0.0172 | 0.0006 |
| $^1D_2$ | 0.1645 | 0.1766 | 0.0108 | –0.0246 |
| | | | | |
| 25.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
| | 56 | 49 | 62.72 | 60.35 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 48.7039 | 48.8364 | 0.1332 | 0.0202 |
| $^3P_0$ | 8.5542 | 8.4968 | 0.6057 | –0.2082 |
| $^3P_1$ | –4.8891 | –4.7272 | 0.1964 | –0.0341 |
| $^3P_2$ | 2.4749 | 2.3712 | 0.1425 | 0.0280 |
| $^1D_2$ | 0.6898 | 0.7665 | 0.0922 | –0.0209 |
| $\varepsilon_2$ | –0.8089 | –0.9582 | 0.1303 | 0.0075 |

| 50.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
|---|---|---|---|---|
| | 218 | 208 | 210.10 | 206.98 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 39.1216 | 39.0859 | 0.1034 | 0.0563 |
| $^3P_0$ | 11.4412 | 11.2540 | 0.3031 | –0.2510 |
| $^3P_1$ | –8.2654 | –8.2992 | 0.0480 | –0.0289 |
| $^3P_2$ | 5.8231 | 5.8834 | 0.0629 | 0.0566 |
| $^1D_2$ | 1.6951 | 1.7153 | 0.0141 | –0.0197 |
| $\varepsilon_2$ | –1.7088 | –1.7186 | 0.0221 | 0.0102 |

| 100.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
|---|---|---|---|---|
| | 155 | 143 | 171.64 | 151.25 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 24.915 | 24.722 | 0.479 | 0.104 |
| $^3P_0$ | 9.481 | 10.690 | 1.197 | –0.276 |
| $^3P_1$ | –13.287 | –13.429 | 0.236 | –0.018 |
| $^3P_2$ | 10.987 | 10.736 | 0.207 | 0.095 |
| $^1D_2$ | 3.794 | 3.630 | 0.100 | –0.021 |
| $\varepsilon_2$ | –2.650 | –2.562 | 0.105 | 0.012 |
| $^3F_2$ | 0.817 | 1.083 | 0.111 | –0.107 |

| 150.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
|---|---|---|---|---|
| | 323 | 309 | 367.34 | 359.09 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 14.167 | 14.306 | 0.374 | 0.137 |
| $^3P_0$ | 4.720 | 4.966 | 0.275 | –0.273 |
| $^3P_1$ | –17.517 | –17.626 | 0.094 | –0.007 |
| $^3P_2$ | 13.990 | 13.943 | 0.069 | 0.122 |
| $^1D_2$ | 5.687 | 5.691 | 0.073 | –0.024 |
| $\varepsilon_2$ | –2.859 | –2.854 | 0.045 | 0.013 |
| $^3F_2$ | 1.200 | 1.010 | 0.089 | –0.127 |
| $^3F_3$ | –2.101 | –1.881 | 0.118 | –0.021 |

| 215.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
|---|---|---|---|---|
| | 197 | 184 | 266.88 | 253.00 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | 4.107 | 4.205 | 0.311 | 0.168 |
| $^3P_0$ | –1.903 | –1.810 | 0.395 | –0.258 |
| $^3P_1$ | –22.311 | –22.152 | 0.204 | 0.005 |
| $^3P_2$ | 16.023 | 15.846 | 0.120 | 0.145 |
| $^1D_2$ | 7.607 | 7.766 | 0.115 | –0.029 |
| $\varepsilon_2$ | –2.685 | –2.550 | 0.085 | 0.012 |
| $^3F_2$ | 1.466 | 1.183 | 0.168 | –0.146 |
| $^3F_3$ | –2.576 | –2.455 | 0.109 | –0.022 |

| | | | | |
|---|---|---|---|---|
| $^3F_4$ | 1.926 | 1.859 | 0.102 | 0.066 |
| $^1G_4$ | 1.050 | 0.986 | 0.048 | –0.011 |
| $\varepsilon_4$ | –1.174 | –1.162 | 0.058 | 0.007 |
| $^3H_4$ | 0.354 | 0.292 | 0.069 | –0.084 |
| | | | | |
| 320.0 MeV | $N_{\rm obs}$ | $N_{\rm df}$ | $\chi^2_{\rm ME}$ | $\chi^2_{\rm SE}$ |
| | 350 | 336 | 357.78 | 353.18 |
| Phase | ME | SE | Error(SE) | $\delta^C_{C+EM}$ |
| $^1S_0$ | –8.064 | –8.142 | 0.349 | 0.200 |
| $^3P_0$ | –12.364 | –12.495 | 0.468 | –0.228 |
| $^3P_1$ | –29.107 | –29.292 | 0.292 | 0.019 |
| $^3P_2$ | 17.368 | 17.508 | 0.133 | 0.166 |
| $^1D_2$ | 9.879 | 9.871 | 0.115 | –0.035 |
| $\varepsilon_2$ | –2.290 | –2.410 | 0.094 | 0.009 |
| $^3F_2$ | 1.257 | 1.321 | 0.160 | –0.168 |
| $^3F_3$ | –2.883 | –2.919 | 0.113 | –0.023 |
| $^3F_4$ | 3.094 | 3.056 | 0.091 | 0.082 |
| $^1G_4$ | 1.553 | 1.565 | 0.065 | –0.012 |
| $\varepsilon_4$ | –1.523 | –1.557 | 0.051 | 0.007 |
| $^3H_4$ | 0.561 | 0.674 | 0.079 | –0.100 |

## FIGURES

FIG. 1. Phase shifts in degrees vs $T_{\rm lab}$ in MeV. Errors are shown only if they are large enough to be plotted. Solid line: multienergy analysis; dashed line: Bystricky *et al.* (Ref. [6]). •: single-energy analyses; ∘: Arndt *et al.* (Ref. [7]); □: Dubois *et al.* (Ref. [18]); ◇: Bugg *et al.* (Ref. [19]).

FIG. 2. Phase-shift combinations of $^3P$ and $^3F$ waves in degrees vs $T_{\rm lab}$ in MeV. Errors are shown for our SE results (bullets) only. Curves and symbols have the same meaning as in Fig. 1.